



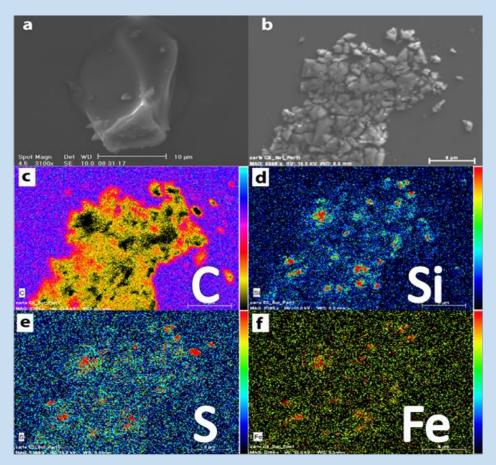


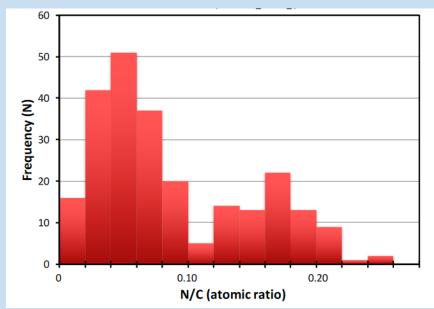
# Effects of Galactic Cosmic Rays on the surface of icy objects, implication for ultra-carbonaceous Antarctic micrometeorites

**Basile AUGE** 

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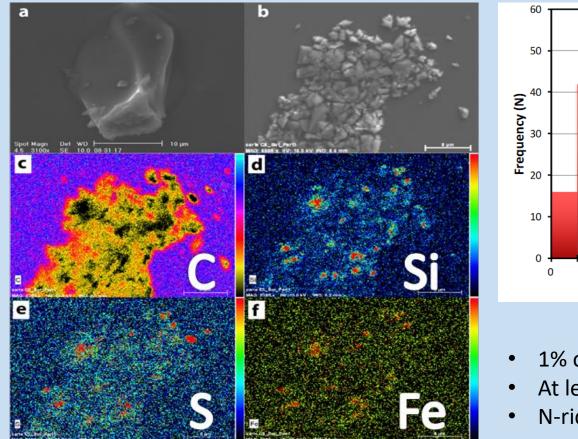
#### Understanding the origin of UCAMM organic matter

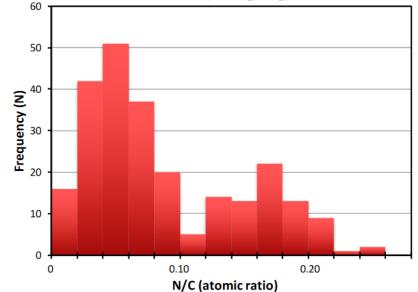




- 1% of Concordia Collection
- At least 50%w of C
- N-rich organic matter (N/C ~12%)

#### Understanding the origin of UCAMM organic matter





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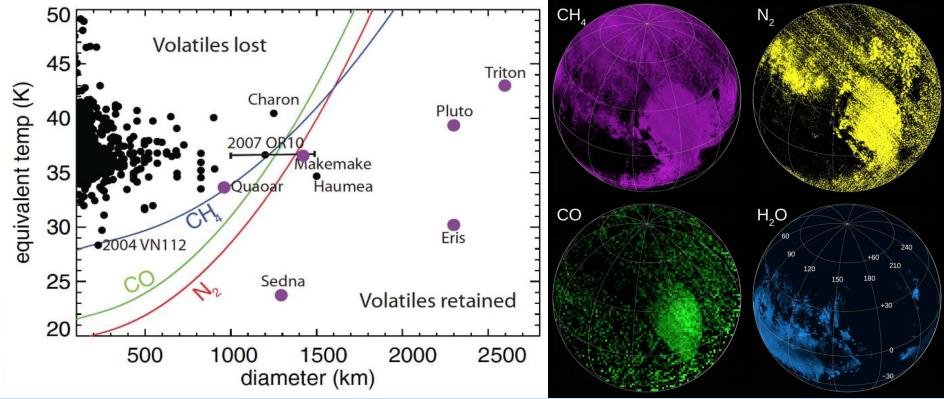
How to produce a mineral depleted N-rich organic matter in the solar system ?

Dartois et al. *Icarus*, 224, 2013 Engrand et al. *LPSC*, 49, 2018

#### N-rich ice surfaces





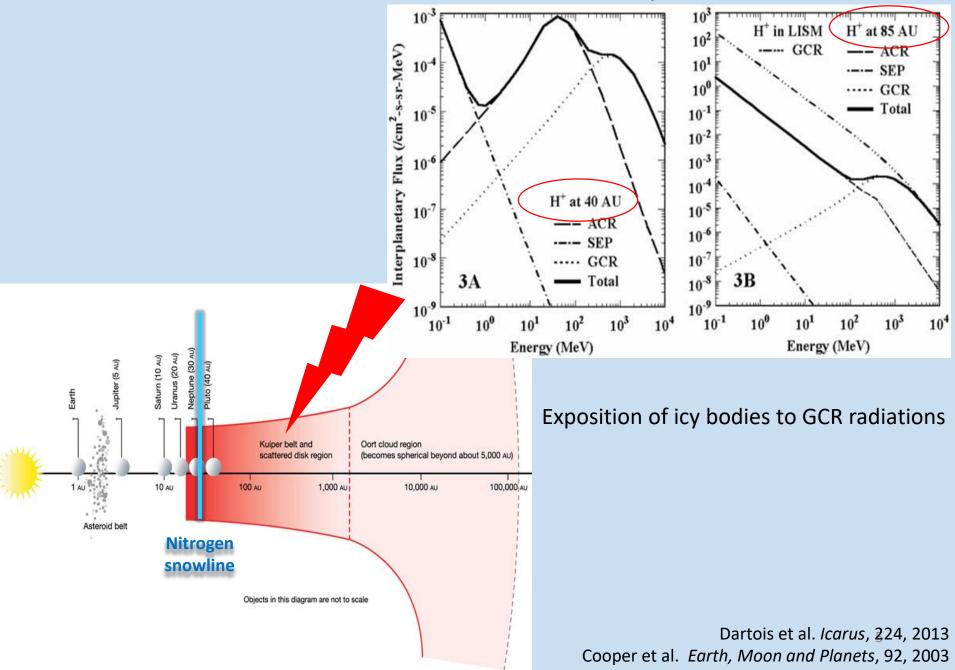


#### Pluto by New Horizons

Kuiper belt – Oort cloud objects can be covered with  $N_2$ -CH<sub>4</sub> ices > 30 AU > 5000 AU

Grundy et al. *Science*, 251, 2016 Brown et al. *ApJ*, 738, 2011

#### Radiative environment : Galactic Cosmic Rays

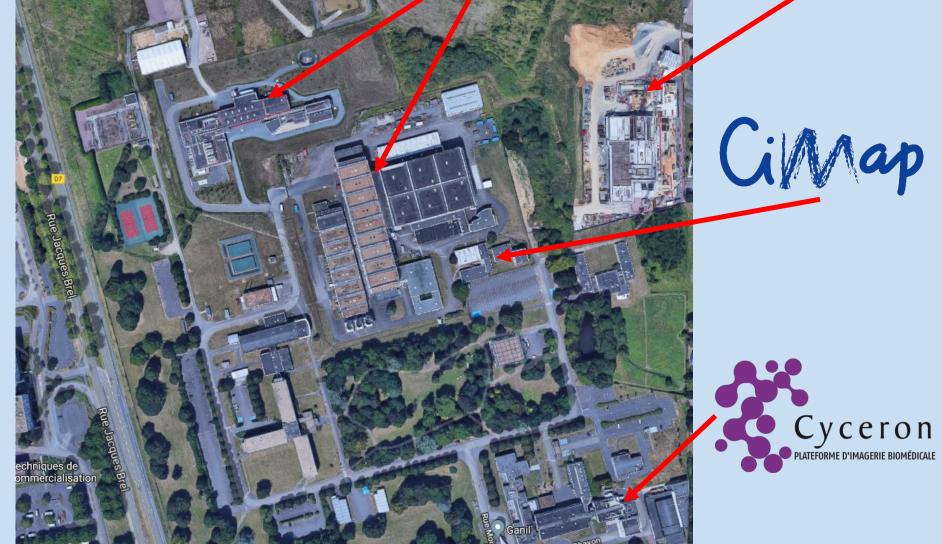


## GANIL (Caen, France)

Large Heavy Ion National Accelerator



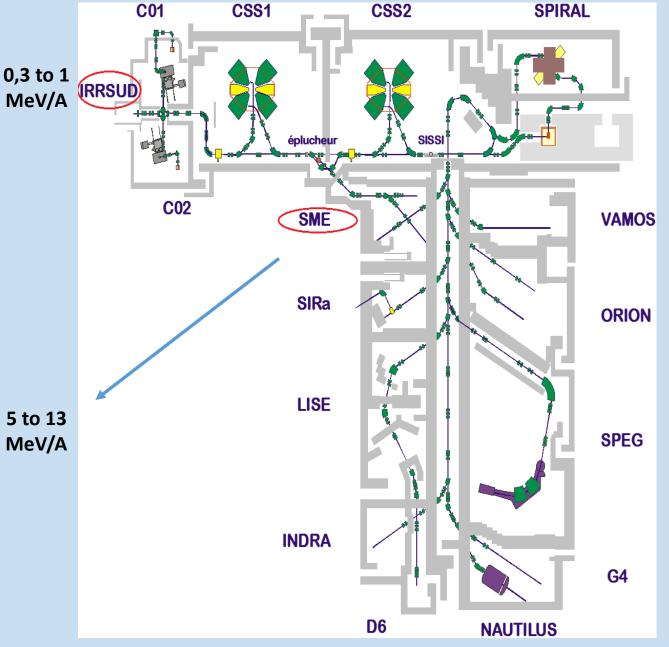




#### Irradiation beam lines







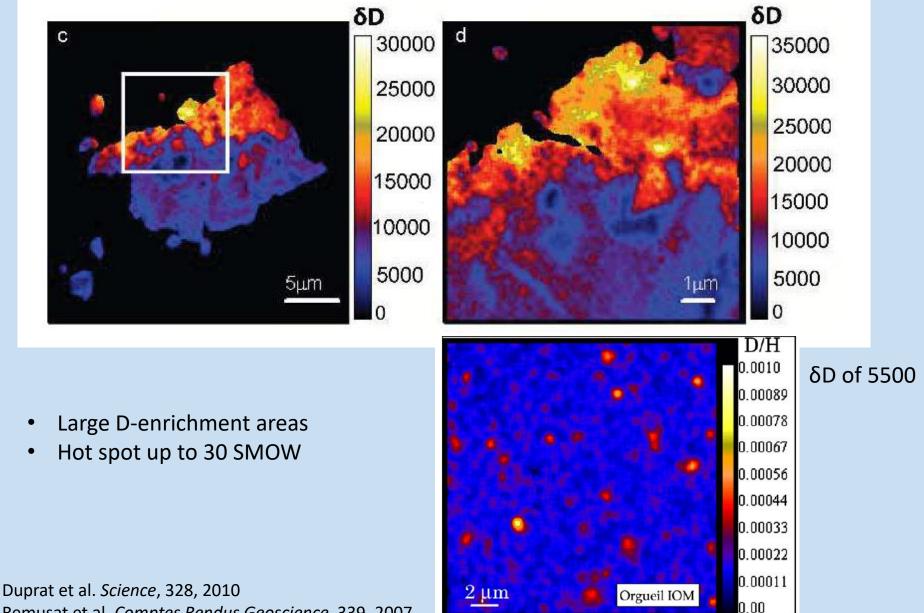
Understanding the origin of UCAMM organic matter

Specific organic matter found in UCAMM can be formed on N-rich icy parent bodies in the outer solar system exposed to GCR

#### A poly-HCN like material can be a precursor for this N-rich O-poor organic matter

Augé et al. Astronomy & Astrophysics, 592, 2016

### A D-rich organic matter

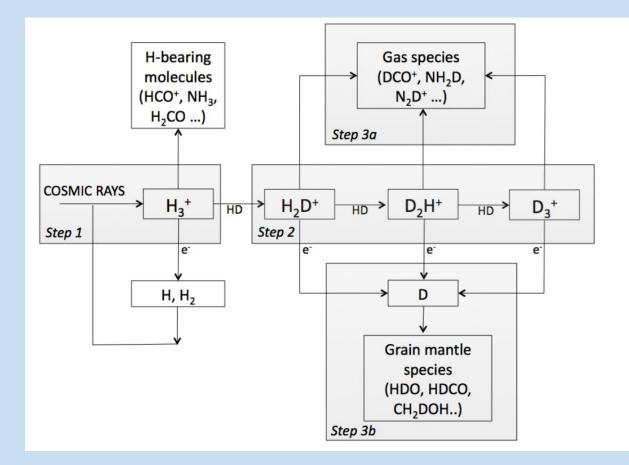


Remusat et al. Comptes Rendus Geoscience, 339, 2007

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#### Hypothesis commonly considered



Ion-molecule reactions at low temperature in the dense phase of a molecular cloud or in a protoplanetary disk

Aikawa et al. *Astronomy & Astrophysics*, 386, 2002 Geiss et al. *Astronomy & Astrophysics*, 93, 1981

#### Secondary processes : irradiation effects

THE ASTROPHYSICAL JOURNAL, 840:35 (11pp), 2017 May 1 © 2017. The American Astronomical Society. All rights reserved. https://doi.org/10.3847/1538-4357/aa6bfc



hysics,

#### X-Ray-induced Deuterium Enrichment of N-rich Organics in Protoplanetary Disks: An Experimental Investigation Using Synchrotron Light

Lisseth Gavilan<sup>1</sup>, Laurent Remusat<sup>2</sup>, Mathieu Roskosz<sup>2</sup>, Horia Popescu<sup>3</sup>, Nicolas Jaouen<sup>3</sup>, Christophe Sandt<sup>4</sup>, Cornelia Jäger<sup>5</sup>, Thomas Henning<sup>6</sup>, Alexandre Simionovici<sup>7</sup>, Jean Louis Lemaire<sup>8</sup>, Denis Mangin<sup>9</sup>, and Nathalie Carrasco<sup>1</sup> <sup>1</sup>LATMOS, Université Versailles St Quentin, UPMC Université Paris 06, CNRS, 11 blvd d'Alembert, F-78280 Guyancourt, France; lisseth.gavilan@latmos.ipsl.fr <sup>2</sup>IMPMC, CNRS UMR 7590; Sorbonne Universités, UPMC Université Paris 06; IRD, Muséum National d'Histoire Naturelle, CP 52, 57 rue Cuvier, Paris F-75231, France <sup>3</sup>SEXTANTS beamline, SOLEIL synchrotron, L'Orme des Merisiers, F-91190 Saint-Aubin, France

<sup>5</sup> Laborator



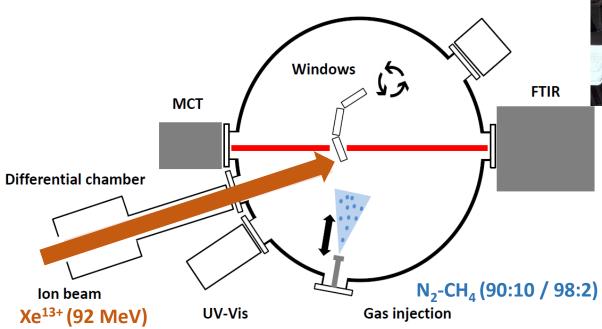
Geochimica et Cosmochimica Acta

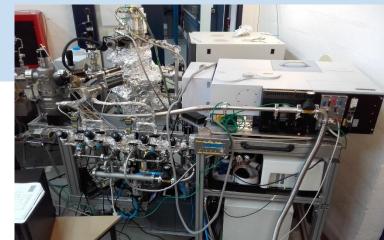
Volume 74, Issue 15, 1 August 2010, Pages 4454-4470

Isotopic anomalies in organic nanoglobules from Comet 81P/Wild 2: Comparison to Murchison nanoglobules and isotopic anomalies induced in terrestrial organics by electron irradiation

Bradley T. De Gregorio <sup>a, b</sup> A ⊠, Rhonda M. Stroud <sup>a</sup> ⊠, Larry R. Nittler <sup>c</sup>⊠, Conel M.O'D. Alexander <sup>c</sup>⊠, A.L. David Kilcoyne <sup>d</sup> ⊠, Thomas J. Zega <sup>a</sup> ⊠

### Experimental set-up





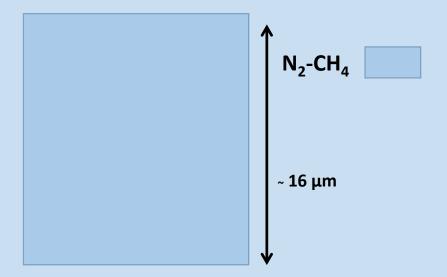


- 1,5 10<sup>-10</sup> mbar at 8 K
- 3 windows
- FTIR and Vis-UV spectrometer

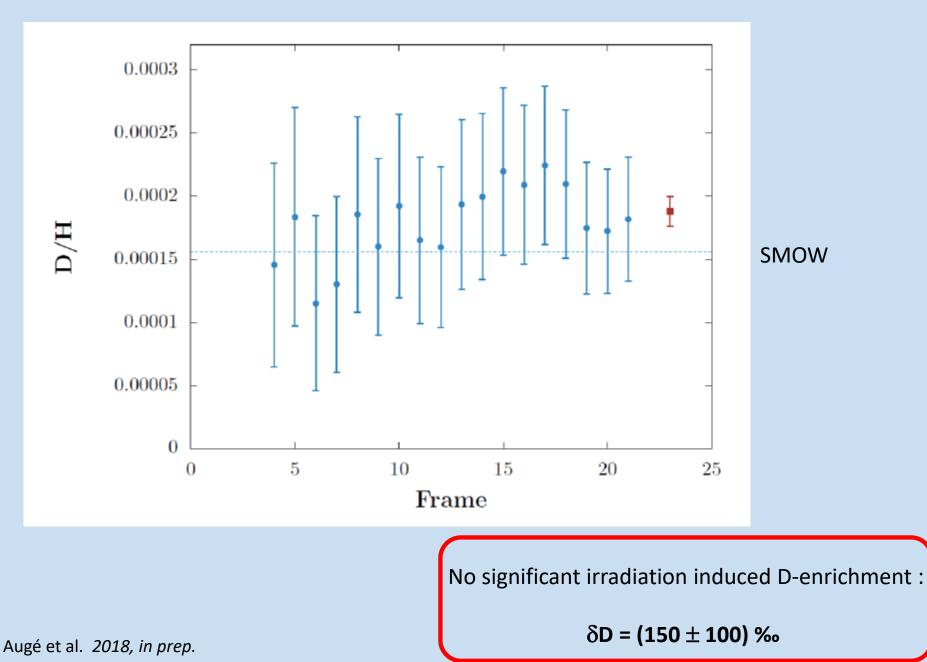
20 mm

#### Irradiation and D-enrichment

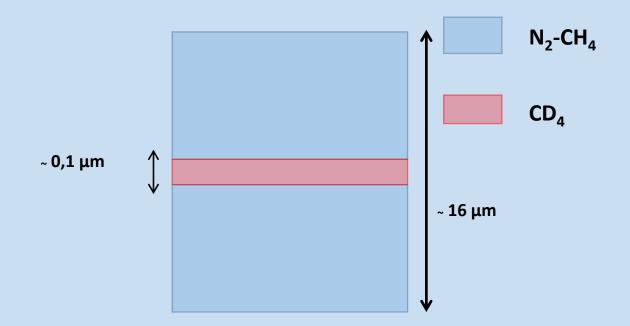
- Irradiation with Xe<sup>13+</sup> (92 MeV) up to 19 eV molec<sup>-1</sup>
- Slow annealing to obtain solid organic residue
- NanoSIMS analyses to obtain D/H ratio or  $\delta D$



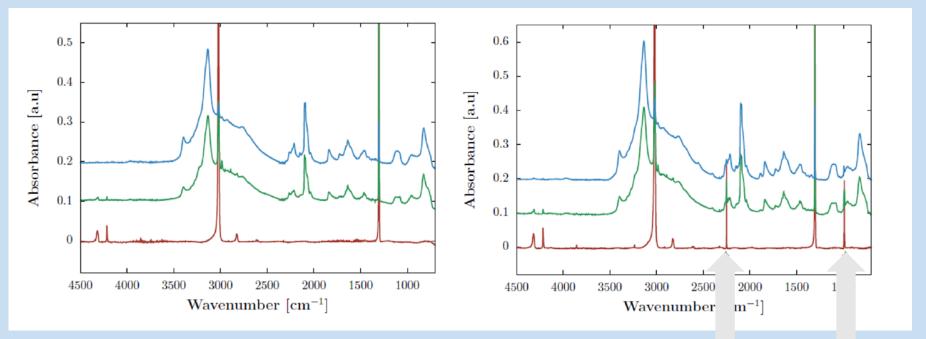
#### Irradiation and D-enrichment



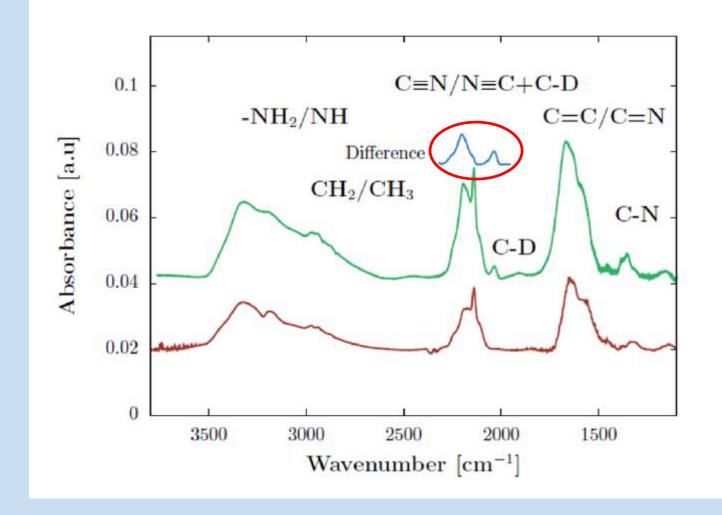
#### Irradiation of a D-rich ices



IR spectra  $N_2$ -CH<sub>4</sub> /  $N_2$ -CH<sub>4</sub> + CD<sub>4</sub>



#### D-rich residue



#### Conclusion

- SHI irradiation does not induce D-enrichment
- Irradiation of D-rich ice induces a D-incorporation in the solid residue in a reasonable astrophysical time scale
- Irradiation keeps the isotopic memory of initial ices

## The scenario can explain both N-rich and O-poor AND Drich organic matter found in UCAMMs